

# MODIS Light

## Option for Low-Risk EOS Cost Reduction

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Hughes Santa Barbara Remote Sensing

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The logo consists of a black rounded rectangle containing the word "HUGHES" in large, bold, white capital letters, and the word "AIRCRAFT" in smaller, bold, white capital letters directly below it.

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**AIRCRAFT**

# MODIS Light Lowers Cost Minimizes Risks

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- Reduce Spacecraft Resources and Launch Mass
  - Delete less critical subsystems (e.g., SRCA)
  - Reduce Scan ( $\pm 45^\circ$ ) and Aperture (6")
  - Repackage to Cut Mass & Volume
  - Reduced Electronics Lowers Power
- Reduce Instrument Recurring Engineering (RE) and Non-RE (NRE) Costs
  - Maximum use of Existing Flight Subsystems cuts NRE
  - Maximum use of Existing Ground Equipment cuts NRE
  - Streamlined Documentation/Manufacturing Process cuts Both
- High Performance and Low Risk
  - 36 MODIS Bands with same IFOVs/Spectral Definition
  - Solar Diffuser/Blackbody Calibration
  - Minimal Impact to MODIS Processing Algorithms

# MODIS to MODIS Light

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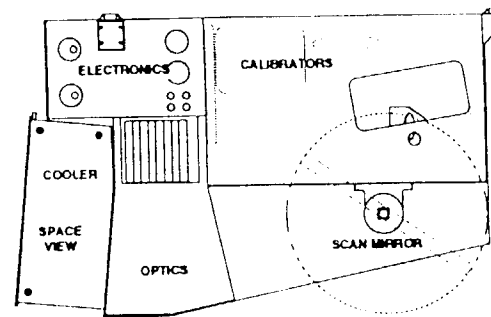
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## Baseline

- Mass 220 kg
- Size 1x1.6x1m
- Power 160W
- Cost No NRE
- Bands 36

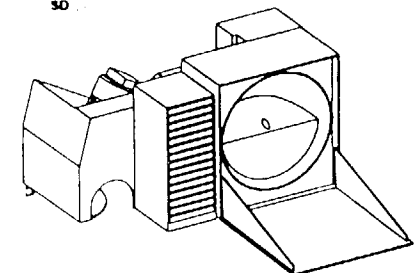
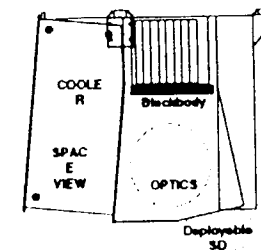
### Calibration

- SRCA
- SDSM
- SD
- BB



## Smaller Scanner Reduced Calibration Smaller Mainframe Repackaged Electronics

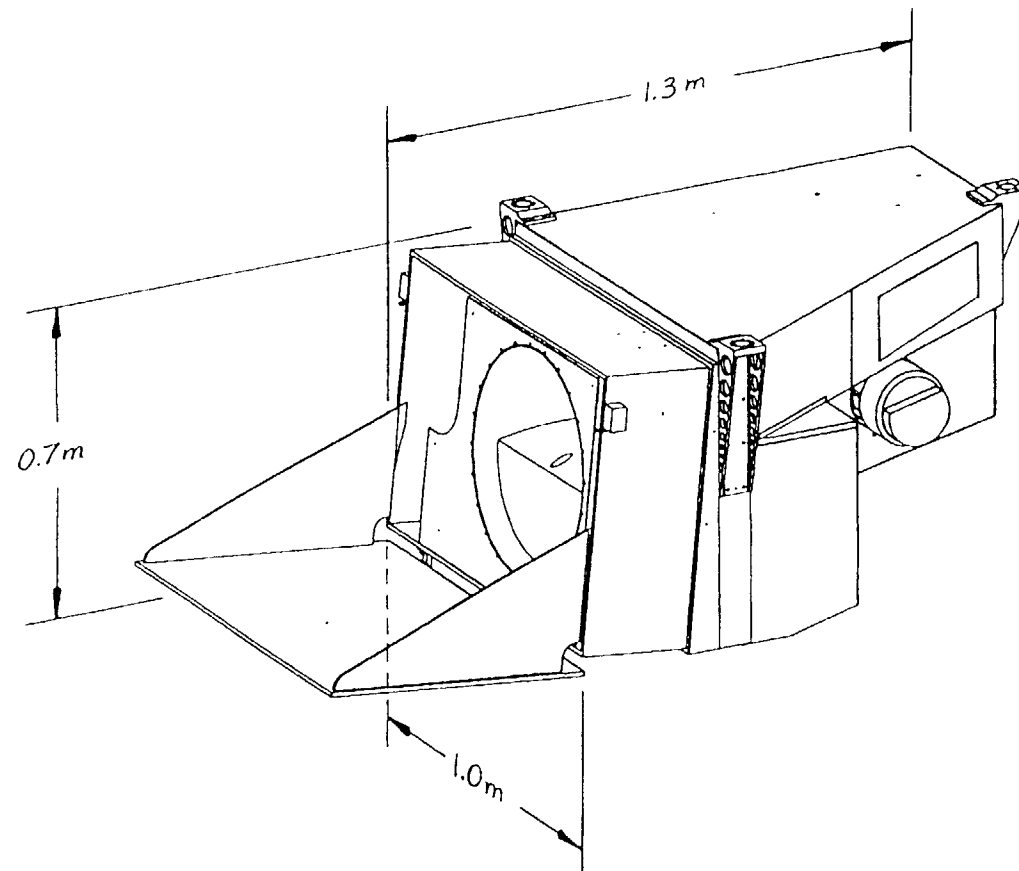
- |          |            |             |
|----------|------------|-------------|
| • Mass   | 100 kg     | Calibration |
| • Volume | 1x0.6x0.7m | • SD        |
| • Power  | 100W       | • BB        |
| • Cost   | Low NRE    |             |
| • Bands  | 36         |             |



# MODIS Light - Baseline Scan Approach

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- Larger than MODIS Light with Tilted Scan
- Baseline Swath Shape and Optical Path

# Design Approach Affects Performance

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<i><b>Parameter</b></i>	<b>Selected MODIS Design</b> (Including MODIS Light)	<b>Other Approaches Considered*</b>
<b>Radiometric</b> Stray Light Calibration  Tunability**	<ul style="list-style-type: none"><li>• Low with 2d Int. Stop</li><li>• Allows Frequent Calibration</li><li>• Small Calibration Target</li><li>• Individual Array per Band</li></ul>	<ul style="list-style-type: none"><li>• Field Stop Limited to 1d</li><li>• Calibration Obstructs Scene; 1/f noise</li><li>• Requires Large, Uniform Target</li><li>• Bands share Detectors: Hi Gain Appls Penalized</li></ul>
<b>Spectral</b> Bandpass Shape Purity	<ul style="list-style-type: none"><li>• Tailored Band-shapes</li><li>• Individual Filters fed by Distributed Blocking</li></ul>	<ul style="list-style-type: none"><li>• Requires Band Synthesis</li><li>• More Difficult to Maintain Constant Center <math>\lambda</math> across FOV; Difficult Out-of-Band Rejection</li></ul>
<b>Spatial</b> Uniformity	<ul style="list-style-type: none"><li>• Full-Field Uniformity</li><li>• Few Detectors Eases Calibration</li></ul>	<ul style="list-style-type: none"><li>• WFOV Edge Distortion</li><li>• More difficult Calibration</li></ul>

\* Narrow/Wide FOV Spectrometers using Gratings, Prisms, or Wedge Filters

\*\*MODIS has multiple similar bands with Application Specific Radiometry

# Summary

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- Straightforward Repackaging Offers Significant Reduction in Mass and Volume
- Reduces Instrument Cost: Low NRE and RE MODIS Miniaturization Approach
- Maximizes MODIS Data Continuity: Minimal Impact to MODIS Processing Algorithms
- Preserves *Optical Bench Assembly* “Heart of MODIS” that Retains MODIS Spectroradiometry